

The Local Group Satellite Galaxies:

Testing Fundamental Physics

Pavel Kroupa
Argelander-Institute for Astronomy
University of Bonn

Pavel Kroupa: AIfA, University of Bonn

- If
- virial equilibrium
 - spherical symmetry
 - isotropic velocity dispersion

then apply *virial theorem* to estimate $\frac{M}{L}$

$$E_{\text{kin}} = -E_{\text{tot}} \quad \longrightarrow \quad \sigma^2 = G \frac{M}{R}$$

$$\text{Luminosity } L = \pi R^2 I \quad \Bigg| \quad I = \text{surf.lumin.}$$

$$\longrightarrow \quad \frac{M}{L} = \frac{1}{G \pi R I} \sigma^2$$

Pavel Kroupa: AIfA, University of Bonn

Background

$$S + d1 + d2 + \dots + d_n$$

$$n \approx 30, \quad d \leq 250 \text{ kpc}$$

Pavel Kroupa: AIfA, University of Bonn

$$\frac{M}{L} = \frac{1}{G \pi R I} \sigma^2 = \frac{1}{G} \frac{\sigma^2 R}{L}$$

Globular clusters :

$$R \approx 10 \text{ pc}, \quad \sigma \approx 10 \text{ km/s}, \quad L \approx 10^5 L_{\odot}$$

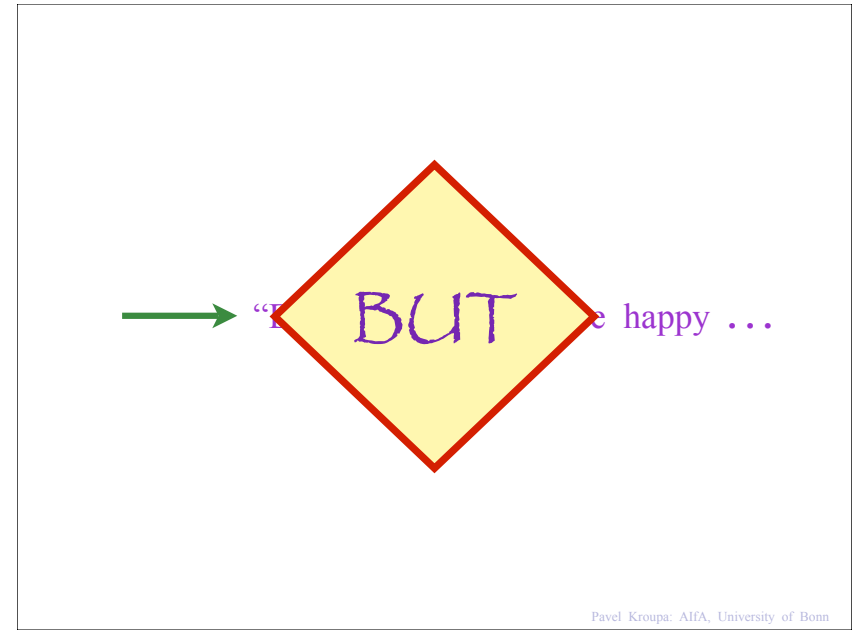
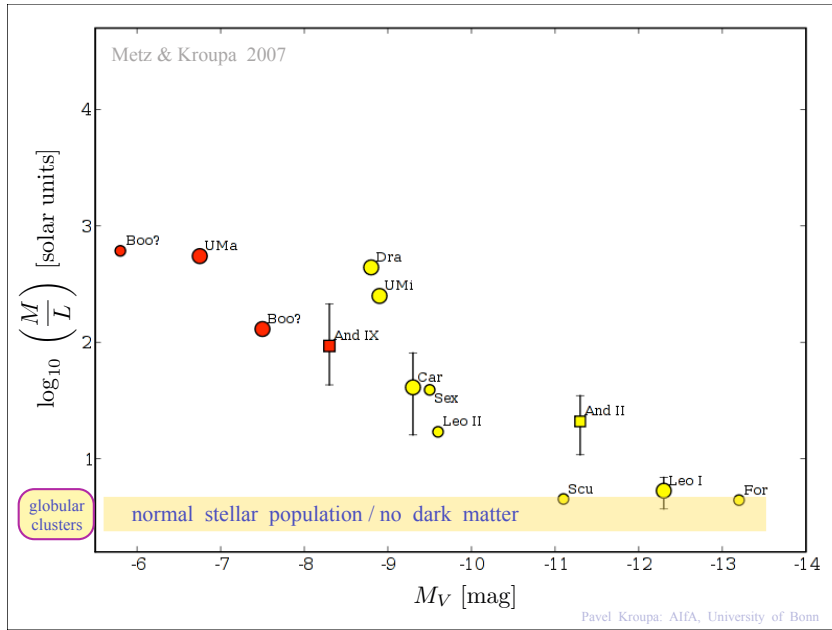
$$\longrightarrow \quad \frac{M}{L} \approx 1$$

Dwarf spheroidal (dSph) satellite galaxies :

$$R \approx 200 \text{ pc}, \quad \sigma \approx 10 \text{ km/s}, \quad L \approx 10^5 L_{\odot}$$

$$\longrightarrow \quad \frac{M}{L} \approx 20$$

Pavel Kroupa: AIfA, University of Bonn



Standard / most-popular interpretation :

The *dwarf-spheroidal* (dSph) satellite galaxies are heavily *dark-matter dominated*;

they are part of the *cosmological sub-structure* surrounding the Milky-Way.

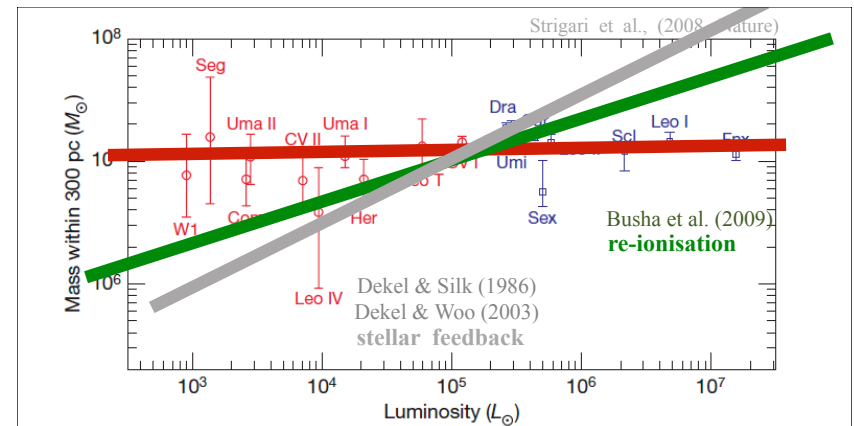


Figure 1 | The integrated mass of the Milky Way dwarf satellites, in units of solar masses, within their inner 0.3 kpc as a function of their total luminosity, in units of solar luminosities. The circle (red) points on the left refer to the newly discovered SDSS satellites, whereas the square (blue) points refer to the classical dwarf satellites discovered pre-SDSS. The error bars reflect the points where the likelihood function falls off to 60.6% of its peak value.

Pavel Kroupa: AIFA, University of Bonn

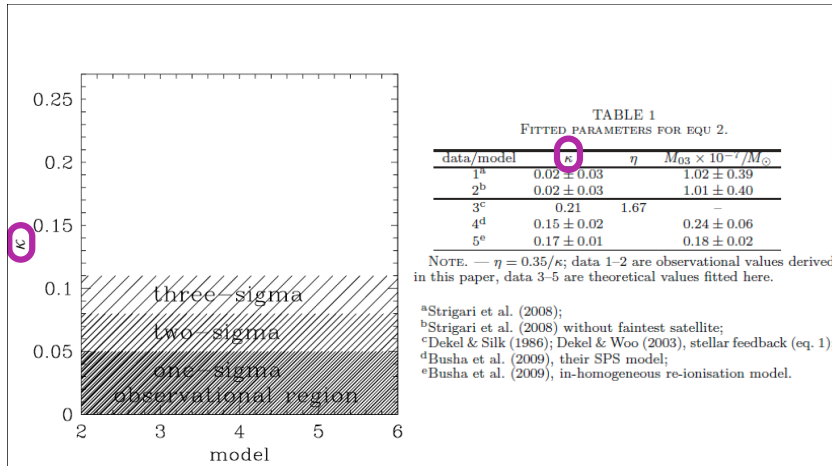


FIG. 1.— The trend κ (eq. 2) for the different models listed in Table 1.

$$\log_{20}(M_{0.3 \text{ kpc}}/M_{\odot}) = \log_{10}(M_{03}/M_{\odot}) + \kappa \log_{10}(L_V/L_{V,\odot})$$

Conservation of energy



it is increasingly difficult to remove sub-parts of bound systems with their increasing binding energy.



The existence of a *theoretical* mass-luminosity relation ($\kappa > 0$)

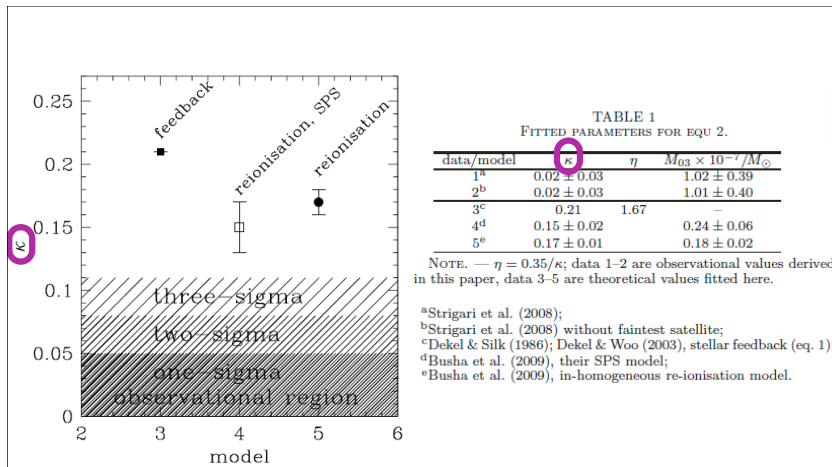


FIG. 1.— The trend κ (eq. 2) for the different models listed in Table 1.

$$\log_{20}(M_{0.3 \text{ kpc}}/M_{\odot}) = \log_{10}(M_{03}/M_{\odot}) + \kappa \log_{10}(L_V/L_{V,\odot})$$

Conservation of energy



it is increasingly difficult to remove sub-parts of bound systems with their increasing binding energy.



The *lack* of an *observed* mass-luminosity relation ($\kappa \approx 0$)

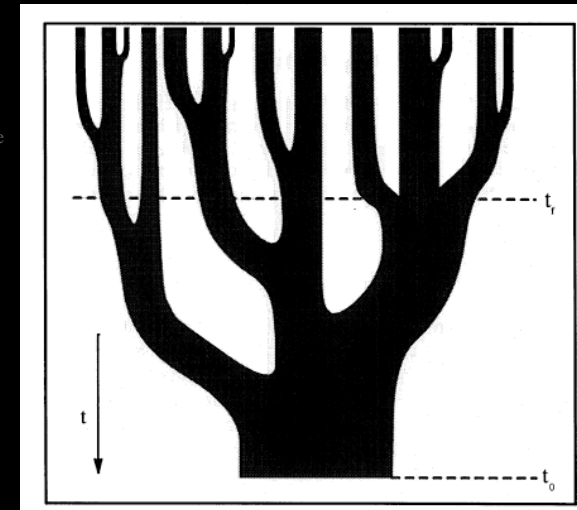
nature apparently does not care about the existence of the putative dark matter halo.

Thus, the *concept*
of dark-matter halos
appears to be
unphysical

Pavel Kroupa: AIfA, University of Bonn

The structure of CDM halos

Lacey & Cole
(1993)



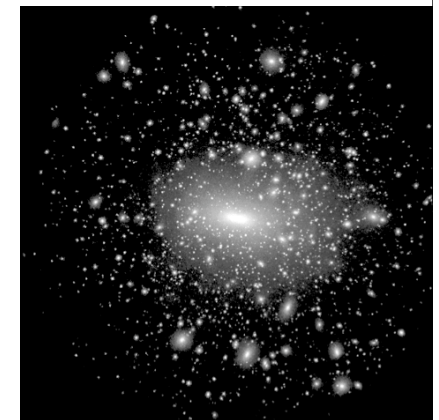
the
beginning

today

Pavel Kroupa: AIfA, University of Bonn

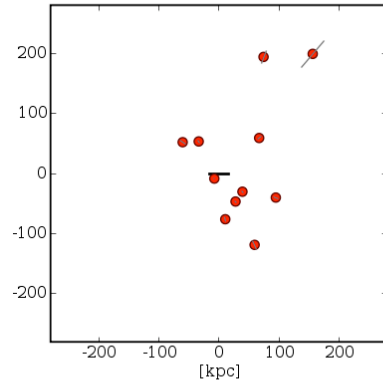
The spatial distribution
of the
MW satellites

Pavel Kroupa: AIfA, University of Bonn



Pavel Kroupa: AIfA, University of Bonn

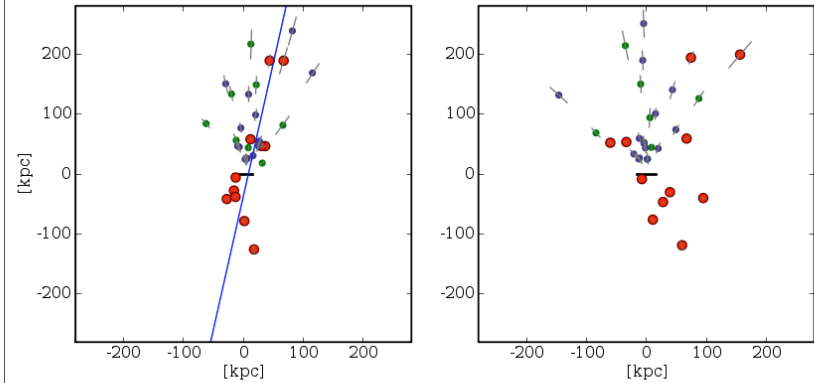
● the 11 “classical” (brightest) satellites



Kroupa, Theis & Boily (2005)
Metz, Kroupa, Jerjen (2006)

MW satellites are in a disk-like configuration:

● the 11 “classical” (brightest) satellites
● new candidate (faint) satellites
● current new candidates (faint) satellites

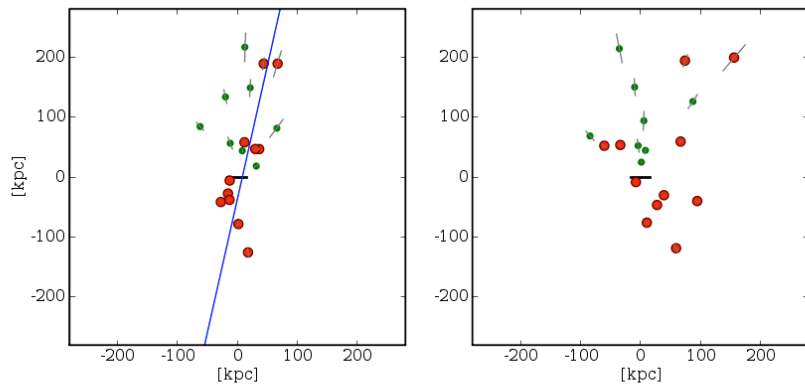


Walsh, Jerjen & Willman (2007)

Pavel Kroupa: AIfA, University of Bonn

MW satellites are in a disk-like configuration:

● the 11 “classical” (brightest) satellites
● new candidate (faint) satellites



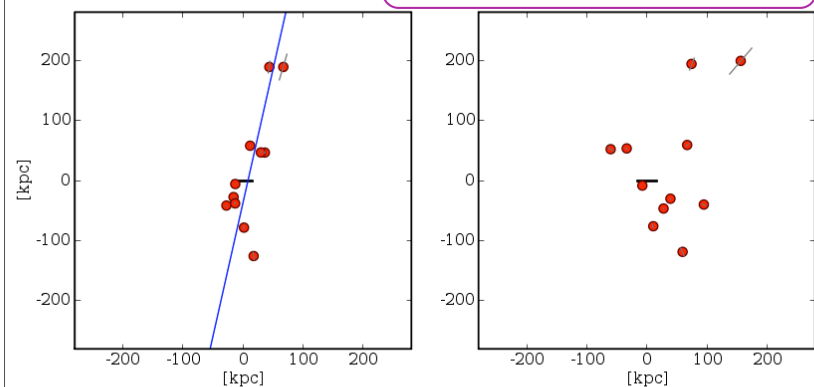
Belokurov et al. (2005);
Zucker et al. (2005)

Pavel Kroupa: AIfA, University of Bonn

Disk of Satellites

➔ *a rotational structure?*

● the 11 “classical” (brightest) satellites

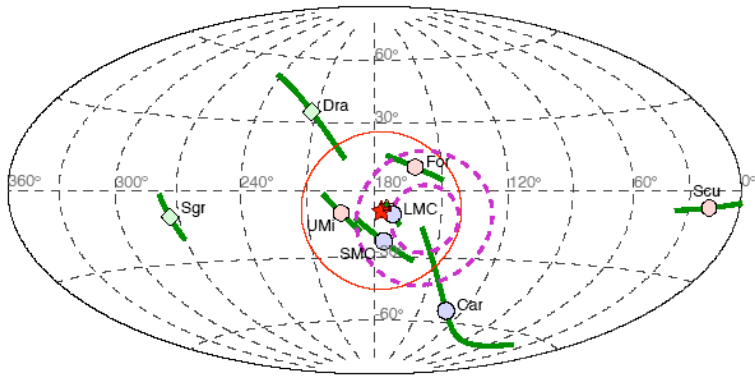


Kroupa, Theis & Boily (2005)
Metz, Kroupa, Jerjen (2006)

Pavel Kroupa: AIfA, University of Bonn

Directions of orbital angular momenta of MW satellites

Metz, Kroupa & Libeskind (2008)



Pavel Kroupa: AIfA, University of Bonn

This *correlated phase-space* population is *inconsistent* with the satellites being dark-matter sub-haloes that fell into the MW halo individually.

Pavel Kroupa: AIfA, University of Bonn



The MW satellite *DoS* is defined mostly by the *outer satellites*; but the *angular momenta* of the *inner satellites* are aligned to the pole of the DoS



the DoS is rotationally supported: the satellites form a *highly correlated phase-space population*

Pavel Kroupa: AIfA, University of Bonn

Andromeda

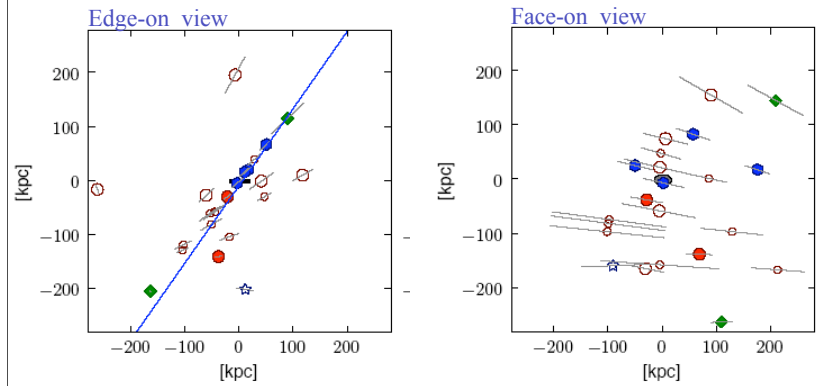


Pavel Kroupa: AIfA, University of Bonn

Andromeda

(Metz, Kroupa, Jerjen 2006)

- dSph
- ◆ dEs, cEs
- dIrrs, dIrr/dSph



Pavel Kroupa: AIfA, University of Bonn

Tidal-dwarf satellite galaxies (TDGs)

(Mirabel, Dottori & Lutz 1992; Duc & Mirabel 1994)

An inherent part of any hierarchical structure formation theory,
and a conservative, *classical approach* to the problem of dSph satellites.

Pavel Kroupa: AIfA, University of Bonn

From α to Ω

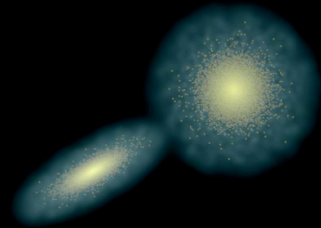
On their origin.

Pavel Kroupa: AIfA, University of Bonn



Pavel Kroupa: AIfA, University of Bonn

Relevance : The collision of two disks at high redshift



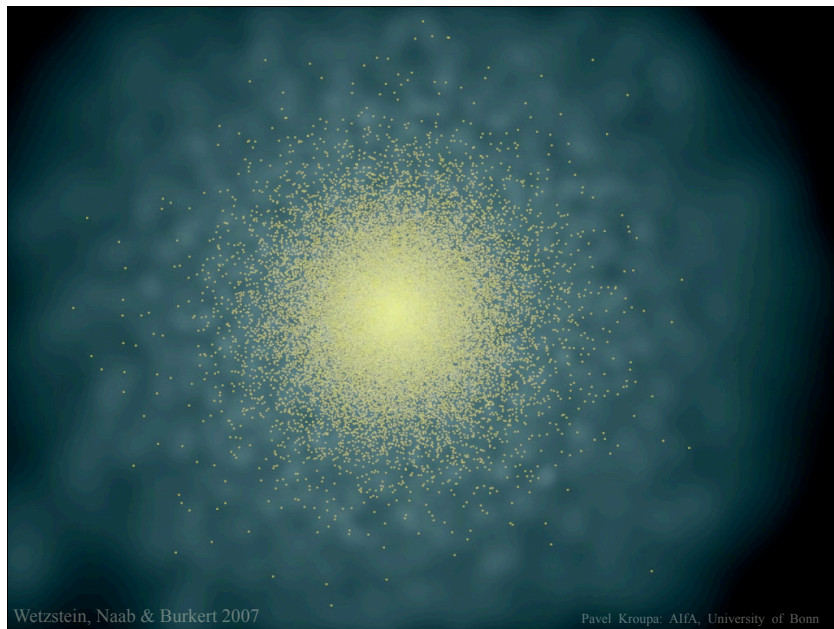
Wetzstein, Naab & Burkert 2007

Pavel Kroupa: AIfA, University of Bonn

TDGs are baryon dominated

(Barnes & Hernquist 1992).

Pavel Kroupa: AIfA, University of Bonn

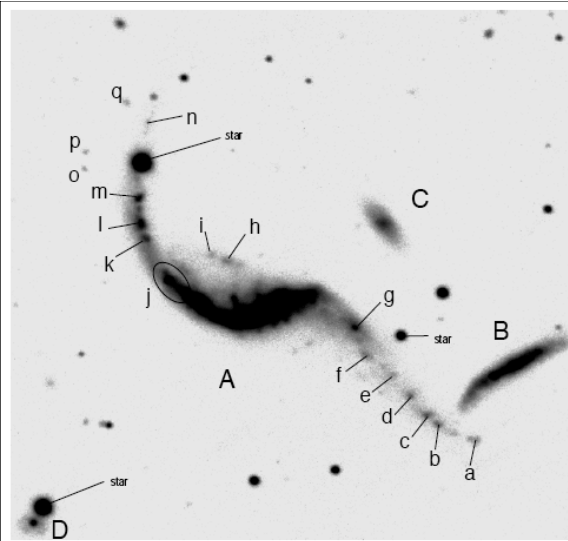


Wetzstein, Naab & Burkert 2007

Pavel Kroupa: AIfA, University of Bonn

An unsettling
Truth ...

Pavel Kroupa: AIfA, University of Bonn



(Weilbacher et al. 2000)

$$N_{\text{TDG}} \approx 14$$

Fig. 21. Identification chart of field 10 around AM 1353-272.

Pavel Kroupa: AIfA, University of Bonn

Okazaki & Taniguchi (2000, ApJ)

“Dwarf galaxy formation induced by galaxy interactions”

Scheme	Collision	Parameter	Result
1	Sp + Sp		S0 + k_1 dEs
2	Sp + S0	a	E + k_2 dEs
		$1 - a$	S0 + S0 + k_3 dEs
3	S0 + S0	b	E + k_4 dEs
		$1 - b$	S0 + S0 + k_5 dEs

NOTE.—In this scheme, a merger between two spiral galaxies evolves not into an elliptical galaxy but into an S0 one. The reason for this is as follows: It is widely accepted that elliptical-like products are formed by dissipationless collapse. Mergers between gas-rich spiral galaxies can achieve a similar physical condition in their final phase. However, if the star formation timescale is longer significantly than the dynamical timescale, the remaining gas will settle to a disk and then the end product will not become an elliptical-like galaxy. This is confirmed by analytical and numerical methods.

Based on the above assumption, we obtain a set of kinetic equations for morphological-type evolution as a consequence of galaxy interactions in the following form:

$$\frac{1}{\gamma} \frac{dn_{\text{Sp}}}{dt} = -2n_{\text{Sp}}^2 - n_{\text{S0}} n_{\text{Sp}}, \quad (1)$$

$$\frac{1}{\gamma} \frac{dn_{\text{S0}}}{dt} = n_{\text{Sp}}^2 + (1 - 2a)n_{\text{Sp}} n_{\text{S0}} - 2bn_{\text{S0}}^2, \quad (2)$$

$$\frac{1}{\gamma} \frac{dn_{\text{E}}}{dt} = bn_{\text{S0}}^2 + an_{\text{S0}} n_{\text{Sp}}, \quad (3)$$

$$\frac{1}{\gamma} \frac{dn_{\text{dE}}}{dt} = k_1 n_{\text{Sp}}^2 + [k_2 a + k_3(1 - a)]n_{\text{S0}} n_{\text{Sp}} + [k_4 b + k_5(1 - b)]n_{\text{S0}}^2, \quad (4)$$

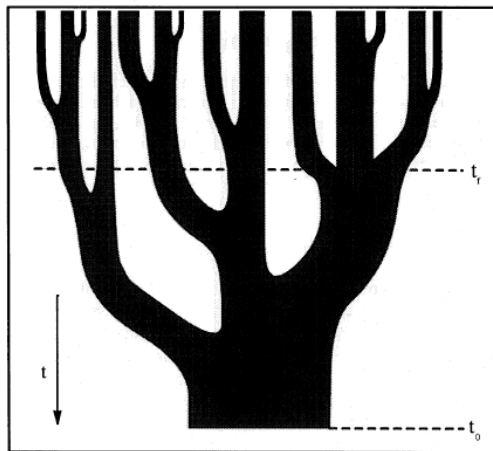
where n_{Sp} , n_{S0} , n_{E} , and n_{dE} are the number densities of spirals, S0s, ellipticals, and dwarfs, respectively, γ is the mean collision rate, and k_i ($i = 1-5$) is the number of dwarfs formed by one collision in each case. Note that the first three equations are the same as those in SN81.

Pavel Kroupa: AIfA, University of Bonn

Okazaki & Taniguchi (2000) :

The galaxy interaction scheme proposed by Silk & Norman (1981)

(Lacey & Cole 1993)



Pavel Kroupa: AIfA, University of Bonn

Okazaki & Taniguchi (2000) :

The galaxy interaction scheme proposed by Silk & Norman (1981)

“can be *responsible for the observed numbers of dEs* in the various environs from poor groups of galaxies to the usual rich clusters of galaxies. The *formation rate of TDGs* is estimated to be **1–2 in each galaxy interaction.**”

i.e. standard cosmology predicts all dE's to be TDGs

But remember, N_{TDG} scales with gas content and thus evolutionary status / cosmological epoch of interacting galaxies (many more formed in the past).

Pavel Kroupa: AIfA, University of Bonn

Within the framework of
standard cosmology,
there is
little room
for
shining cosmological
sub-structures
with $< 10^{10} M_{\odot}$!

(taking account of only the TDGs,
not even counting “fireballs”)

Pavel Kroupa: AIfA, University of Bonn

There are thus
5 essentially
independent
arguments :

Pavel Kroupa: AIfA, University of Bonn

A
contradiction
in *standard cosmological theory*
thus emerges :

theory + observation :
a large fraction (if not all) of
observed $\lesssim 10^{10} M_{\odot}$ sub-structures are TDGs



Previous and current attempts to get the
 $< 10^{10} M_{\odot}$ dark matter subhaloes
to shine would have been ill-fated...

Pavel Kroupa: AIfA, University of Bonn

- 1) *Lack* of a *mass-luminosity relation* (binding energy does not play a role)
and
all haloes having the *same mass* is *incompatible* with the individual
growth histories inherent to the dark-matter hypothesis.
→ dark-matter halo masses appear to be unphysical.
- 2) The satellites form a correlated phase-space population.
This is *incompatible* with them being individual haloes that fell into the
MW halo independently.
→ dark-matter halo masses appear to be unphysical.
- 3) The satellites form a *correlated phase-space population*.
This can only really imply they are *ancient tidal-dwarf galaxies* (TDGs).
Their high M/L ratios
→ dark-matter halo masses appear to be unphysical.

Pavel Kroupa: AIfA, University of Bonn

4) The *number of TDGs* produced in standard cosmology accounts for all dE galaxies.

→ there is not much room for shining dwarf dark matter haloes.

5) *Young TDGs* are observed to have flat rotation curves. But they cannot contain dark matter

→ "dark matter" appears to be a *manifestation of non-Newtonian dynamics*.

The END

The *failure* of the dark-matter ansatz



Newtonian dynamics seems to *fail*
(flat rotation curves and high M/L ratio of dSph satellites).

We therefore have to face the real possibility that
we probably live in a *non-Newtonian universe*.

Popular examples of alternative dynamics theories:
MOND, MOG.

M/L_V vs mass :

Dabringhausen, Hilker & Kroupa 2008

