





NAHUAL:

### A MID & HIGH RESOLUTION IR SPECTROGRAPH FOR A 10-M SEGMENTED TELESCOPE (GTC)

Motivations:

To extend the parameter space of planet searches (ages, masses, rotation).

Follow-up of NIR transit searches.

Very low-mass primaries (Martin et al. 2006, Blake et al. 2007, T Tauri stars (Huélamo et al. 2008), Young MS stars, Red giants

Zapatero Osorio et al. 2007, 2009)



# **Community interest**

- 7 workshops held so far (La Gomera, Segovia, Jena, Cádiz, Miami, Fuerteventura, Sintra).
- Next workshop in Fuerteventura, 16-18 Dec 2009, http://www.reunprep-cons11-fuerteventura.com/
- About 100 different participants in those workshops
- Seed for instrument proposed for Calar Alto 3.5-m (CARMENES)
- Other similar projects in other observatories (PRVS)
- Need for high-precision RV capabilities mentioned in Exoplanet Task Force and Blue Dots reports.
- Funding collected since 2005:
- 600 Keuros from MEC
- 200 Keuros from Tautenburg
- 100 Keuros from IAC
- 40 Keuros in Ireland
- 40 Keuros from Lisbon



# **Observing modes**

NAHUAL team driven: Single object. High stability (no moving parts). R=60,000. Wav. Range=0.9—1.8 microns. FOV=3 arcsec. CD completed.

GTC driven: Multi-object capability. R=20,000. Range=0.95—2.45 microns. FOV≈10 arcmin. CD study funded.

# Long-term NIRSPEC RV data (2001-2008) of VB10 has rms of 300 m/s Zapatero Osorio, Martin, et al. 2009, A&A



Fig. 3.— NIRSPEC radial velocities of vB 10 (filled circles) folded in phase using the same orbital parameters than in Fig. 2 (*left panel*), and using slightly different orbital period, eccentricy, and argument of periapsis (*right panel*). These spectroscopic observations cannot constrain the orbital solution precisely, but they are consistent with the presence of a small body aroud vB 10. The horizontal dotted line denotes the "systemic" velocity of the pair. Note that two periods are depicted in the diagrams.

#### Mass of companion ≈ 6.4 Jupiters; Pravdo & Shaklan 2009

What can we do to improve the precision?



Deconvolution by F. Rodler, precision 120 m/s



### Nahual project web page





### Conceptual mechanical design







# High stability cryostat



#### Collaboration between Arcetri, IAC and Lisbon

### Single vessel e=40 mm (wall thickness)









SUB =8



## Prisms of ZnSe

Ohara GmbH 150.00 ±1.0 mm 50.00 ±1.0 mm 40.000 euros



## Satisfactory delivery to TLS



Zemax model to place in pupil (returning mirror) to consider the effects of The prisms WFE. It contains 37 zerniques (with tilt and focus removed). - There are two passes and is scaled from 0.632 to the working wavelength.





### Gas cells





### Gas cell development

HE



# Mid-resolution gas cell spectra



Collaboration with R. Peale group (UCF)





### NAHUAL SENSITIVITY



### Figure by M.R. Zapatero Osorio

3 m/s ((3 sigma) in 300 s for K=9 and D = 10 m.

D=40m => K=12-14 => L and T dwarfs => lowest mass planets around the lowest mass primaries?

Semi-major axis (AU)



# Summary

- Study to consider NAHUAL as mid and high-resolution nearinfrared spectrograph common-user @ GTC in 2016 approved.
- 5 years of experience and about 1 million euros of development investment as of 2009.
- NAHUAL experience is applicable to other (bigger or smaller) telescopes, such as CARMENES @ Calar Alto and the ELTs.
- Calar Alto status uncertain ...



# Crazy things happen

