

## The Rapid Imager for Surveys of Exoplanets

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Outline

Scientific Motivation - TTV

High precision photometry

The Liverpool Telescope

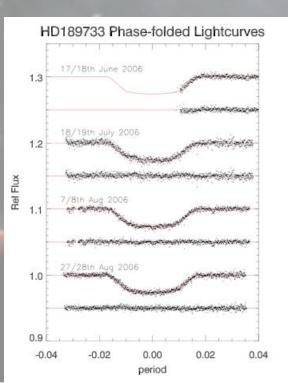


• TTV and other science results

• The future

## A short cut to *Farth mass* objects?

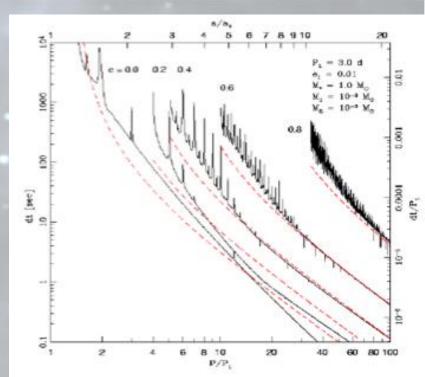
Holman and Murray (2005) – TTV from Jupiter mass companion in >1 yr orbit or hot Earth (P~40-50d) detectable.

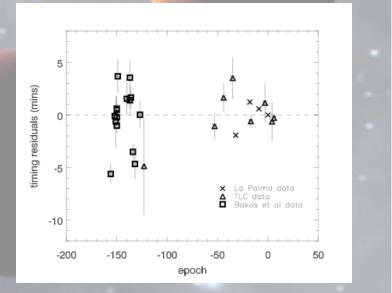


First attempt

NOT/LT/ WHT, Δt ~ 5-15 sec







### How to do high precision photometry

- Sort the Telescope! - Doughnuts....



- If you take flats think about what you doing
- Autoguiding

- Systematic noise sources (filter wheels, temp changes, colour terms etc)

### Liverpool Telescope



2.0m RC optics, alt-az mount. Full robotic control

Optimized for ESP photometric work



## RISE at the T: why we built



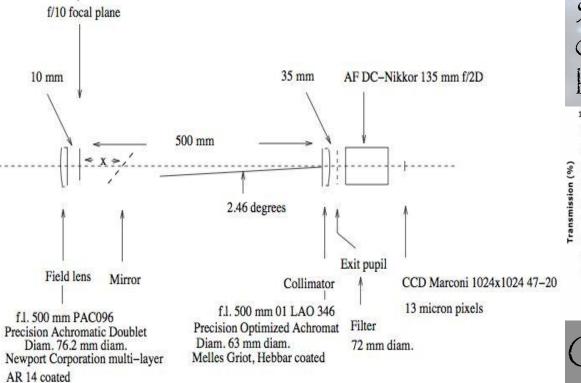
RISE: Optical - John Meaburn, Mechanical - JMU, Software + Construction - QUB

Wide-field, focal reducer for the LT. Uses a frame transfer e2v sensor.

Designed for "rapid" continuous photometry of bright stars with plenty of comparison objects

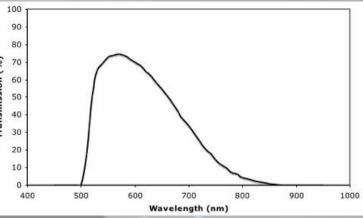
## RISE Characteristics

#### FOV 43 mm and 7.4 arcmin diam. CCD pixel size 0.48 arcsec



#### 9.2 arcmin field 0.54"/pixel. (Jsually used binned 2x2

9.2 arcmin

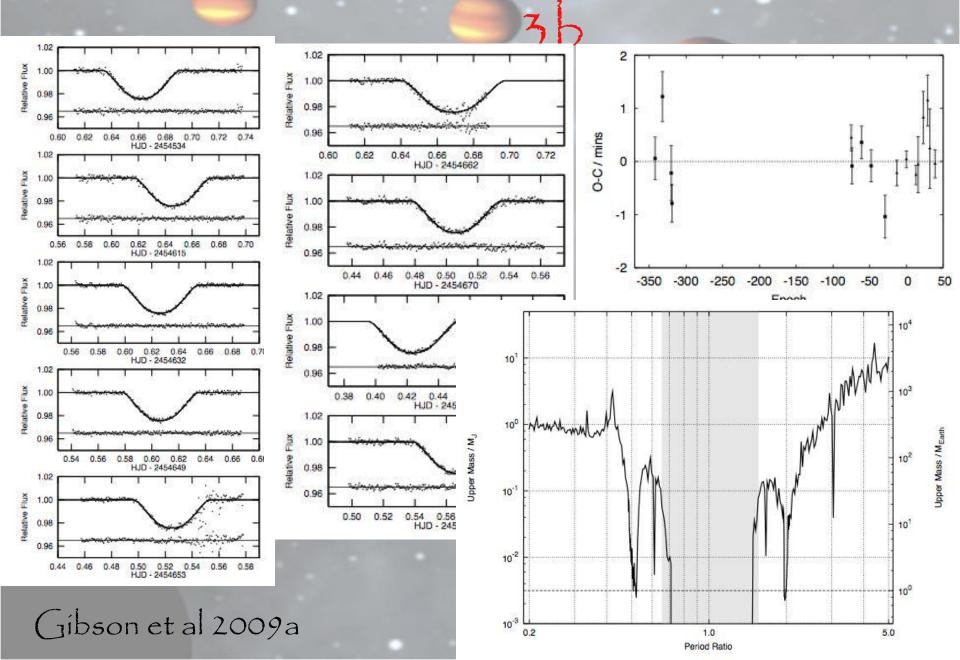


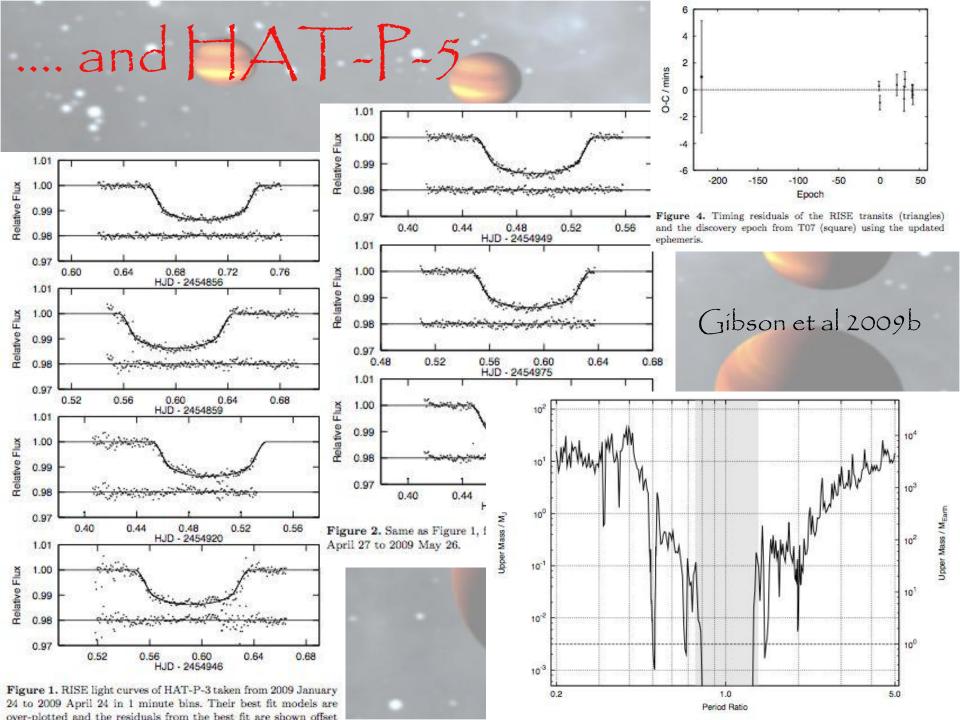
CCD: gain 2.3 e/adu, 10 e ron. Fixed V+R filter

Optical layout matched to LT

Minimum exposure 0.8s, no readout time

#### V observations of



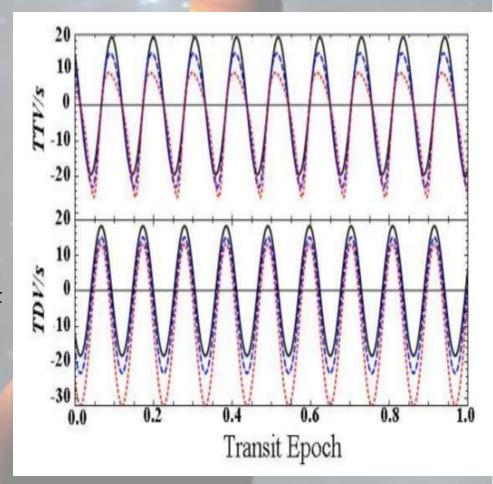


### Transit Duration Variations

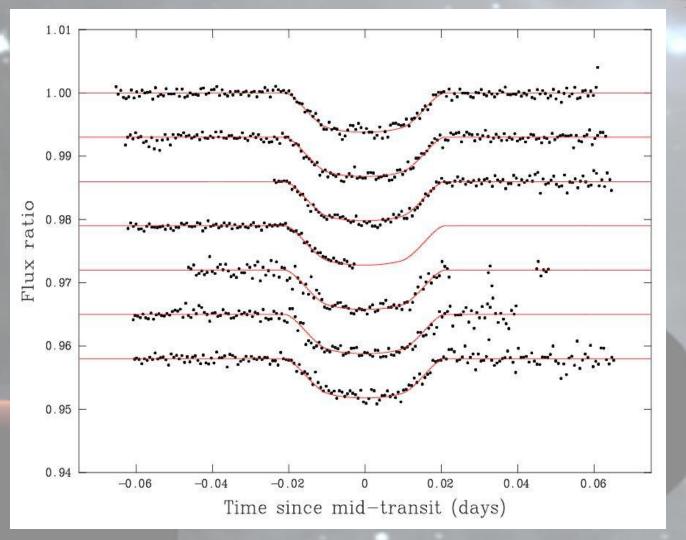
Kipping 2009 Consider an Exo-moon:

- Planet/moon orbit barycentre => oscillatory motion for planet in velocity and position during exo-moon orbit => TTV in usual way but also change in transit duration

- Consider geometry: TTV + TDV out of phase by 90 deg => powerful diagnostic.



# LT/RISE observations of GJ436



Fossey, Kipping etc Feb-june 2009 ΔTTV ~ 14.1s

#### ΔTDV ~13.5s

Extremely challenging but significant and just do-able with RISE (first 4 lc's have errors of 15sec).

## Using RISE in RoPACS

In general the RoPACS targets will be relatively faint at about  $R \sim 15$ , however, exposures of 30-60 sec will still deliver photometry at the millimag level => superearths (early M dwarfs). Several things will be important:

- 1) Calibration files obtained specifically for the night of observations (this isnt normally the case!).
- 2) Reasonably accurate ephemeris very important to get 30-45 minutes of pre- and post-transit photometry.
- 3) (Jseful number of comparison stars (preferably of the same colour).

Conclusion: best use of LT is for publication quality lc's and not really for filtering mimics

## Current RISE ESP Proposals/Projects

1) SuperWASP planet discovery light curves (7 published papers 2008-9) 2) TTV continuation with WASP-12b (2 published papers in 2009) 3) TDV/TTVGJ436 continuation (Fossey/Kipping etc) 4) HAT-P-13 - first double planet - predicted TTV ~20sec, can be used to derive inclination of non eclipsing object with great accuracy. 5) WYFCAM Transit Survey M-dwarf candidates (Pinfield/Hodgkin etc)

6) ESP host star rotation periods

## RISE 2!



Greek 2.3m Arístarchos Telescope.

Original idea from John Meaburn - Collaboration of original partners - Need a JMU science collaborator interested in ESP observational science.

Slightly improved copy of LT/RISE. MoU gives 3n/month for 3yr, more possible

## Conclusions

The LT has been optimized for obtaining this kind of photometry

RISE is commissioned and is taking (generally) extremely high quality data of ESP targets

RISE/LT on RoPACS targets would be able to deliver publication quality lc for objects as small as SuperEarths (early M dwarfs).